

CHAPTER 6

TELESCOPES: PORTALS OF DISCOVERY

6.1 Eyes and Cameras: Everyday Light Sensors

Our goals for learning:

- How does your eye form an image?
- How do we record images?

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Refraction

- Refraction is the bending of light when it passes from one substance into another
- Your eye uses refraction to focus light

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Focusing Light

- Refraction can cause parallel light rays to converge to a focus

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Image Formation

- The focal plane is where light from different directions comes into focus
- The image behind a single (convex) lens is actually upside-down!

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Focusing Light

Digital cameras
detect light with
charge-coupled
devices (CCDs)

- A camera focuses light like an eye and captures the image with a detector
- The CCD detectors in digital cameras are similar to those used in modern telescopes

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What have we learned?

- ⦿ How does your eye form an image?
 - It uses refraction to bend parallel light rays so that they form an image.
 - The image is in focus if the focal plane is at the retina.
- ⦿ How do we record images?
 - Cameras focus light like your eye and record the image with a detector.
 - The detectors (CCDs) in digital cameras are like those used on modern telescopes

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6.2 Telescopes: Giant Eyes

Our goals for learning:

- ⦿ What are the two most important properties of a telescope?
- ⦿ What are the two basic designs of telescopes?
- ⦿ What do astronomers do with telescopes?

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What are the two most important properties of a telescope?

1. **Light-collecting area:** Telescopes with a larger collecting area can gather a greater amount of light in a shorter time.
2. **Angular resolution:** Telescopes that are larger are capable of taking images with greater detail.

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Angular Resolution

- The *minimum* angular separation that the telescope can distinguish.
- Ultimate limit to resolution comes from interference of light waves within a telescope.
- Larger telescopes are capable of greater resolution because there's less interference

Interactive Figure

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Angular Resolution

- The rings in this image of a star come from interference of light wave.
- This limit on angular resolution is known as the **diffraction limit**

Close-up of a star from the Hubble Space Telescope

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What are the two basic designs of telescopes?

- ◉ **Refracting telescope:** Focuses light with lenses
- ◉ **Reflecting telescope:** Focuses light with mirrors

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Refracting Telescope

- ◉ Refracting telescopes need to be very long, with large, heavy lenses

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Reflecting Telescope

- ◉ Reflecting telescopes can have much greater diameters
- ◉ Most modern telescopes are reflectors

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What do astronomers do with telescopes?

- ◉ **Imaging:** Taking pictures of the sky
- ◉ **Spectroscopy:** Breaking light into spectra
- ◉ **Timing:** Measuring how light output varies with time

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Imaging

- Astronomical detectors generally record only one color of light at a time
- Several images must be combined to make full-color pictures

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Imaging

- Astronomical detectors can record forms of light our eyes can't see
- Color is sometimes used to represent different energies of nonvisible light

Interactive Figure

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Spectroscopy

- A spectrograph separates the different wavelengths of light before they hit the detector

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Spectroscopy

- Graphing relative brightness of light at each wavelength shows the details in a spectrum

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Timing

- A light curve represents a series of brightness measurements made over a period of time

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Want to buy your own telescope?

- Buy binoculars first (e.g. 7x35) - you get much more for the same money.
- Ignore magnification (sales pitch!)
- Notice: aperture size, optical quality, portability.
- Consumer research: Astronomy, Sky & Tel, Mercury. Astronomy clubs.

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What have we learned?

- What are the two most important properties of a telescope?
 - Collecting area determines how much light a telescope can gather
 - Angular resolution is the minimum angular separation a telescope can distinguish
- What are the two basic designs of telescopes?
 - Refracting telescopes focus light with lenses
 - Reflecting telescopes focus light with mirrors
 - The vast majority of professional telescopes are reflectors

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What have we learned?

- What do astronomers do with telescopes?
 - Imaging
 - Spectroscopy
 - Timing

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6.3 Telescopes and the Atmosphere

Our goals for learning:

- How does Earth's atmosphere affect ground-based observations?
- Why do we put telescopes into space?

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How does Earth's atmosphere affect ground-based observations?

- The best ground-based sites for astronomical observing are
 - Calm (not too windy)
 - High (less atmosphere to see through)
 - Dark (far from city lights)
 - Dry (few cloudy nights)

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Twinkling and Turbulence

Turbulent air flow in Earth's atmosphere distorts our view, causing stars to appear to twinkle

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What have learned?

- How does Earth's atmosphere affect ground-based observations?
 - Telescope sites are chosen to minimize the problems of light pollution, atmospheric turbulence, and bad weather.
- Why do we put telescopes into space?
 - Forms of light other than radio and visible do not pass through Earth's atmosphere.
 - Also, much sharper images are possible because there is no turbulence.

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6.4 Eyes and Cameras: Everyday Light Sensors

Our goals for learning:

- How can we observe nonvisible light?
- How can multiple telescopes work together?

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Radio Telescopes

- A radio telescope is like a giant mirror that reflects radio waves to a focus

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X-Ray Telescopes

- Focusing of X-rays requires special mirrors
- Mirrors are arranged to focus X-ray photons through grazing bounces off the surface

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Gamma Ray Telescopes

- Gamma ray telescopes also need to be in space
- Focusing gamma rays is extremely difficult

Compton Observatory

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Interferometry

- Interferometry is a technique for linking two or more telescopes so that they have the angular resolution of a single large one

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